

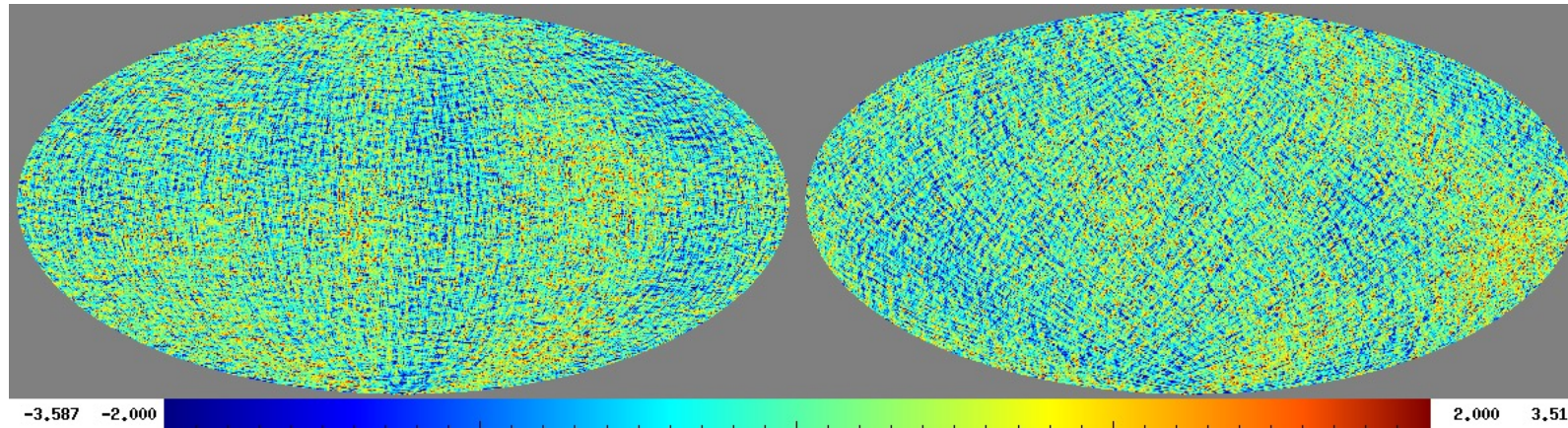
# Pico

Commander updates 13.01.2022  
Ragnhild Aurlien and the Oslo group

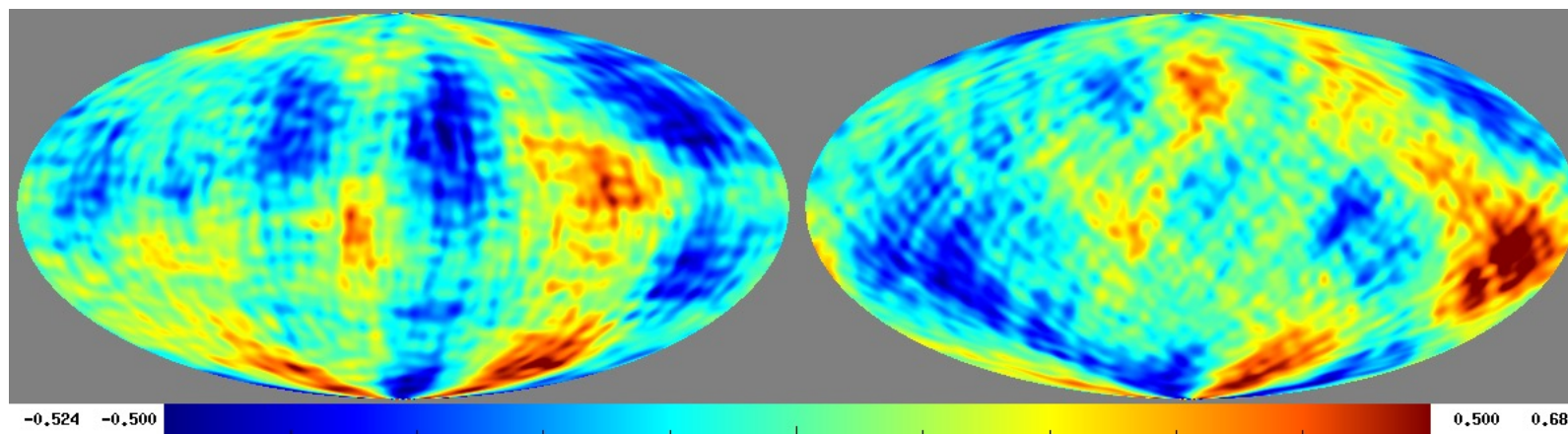
# 90.91

- Running Commander1 fitting 6 parameters pr pixel
  - CMB ( $A_{cmb}$ )
  - One dust model ( $A_{dust}, T_{dust}, \beta_{dust}$ ) (nside 256)
  - Synchrotron model ( $A_{sync}, \beta_{synch}$ ) (nside64)
- 10 simulation sets with  $r = 0$  and 10 with  $r = 0.003$
- Masking 25% of the sky
  - Mask based on the  $\chi^2$  map
- Infilling inside mask:  $\sim 180$  samples pr CMB realization

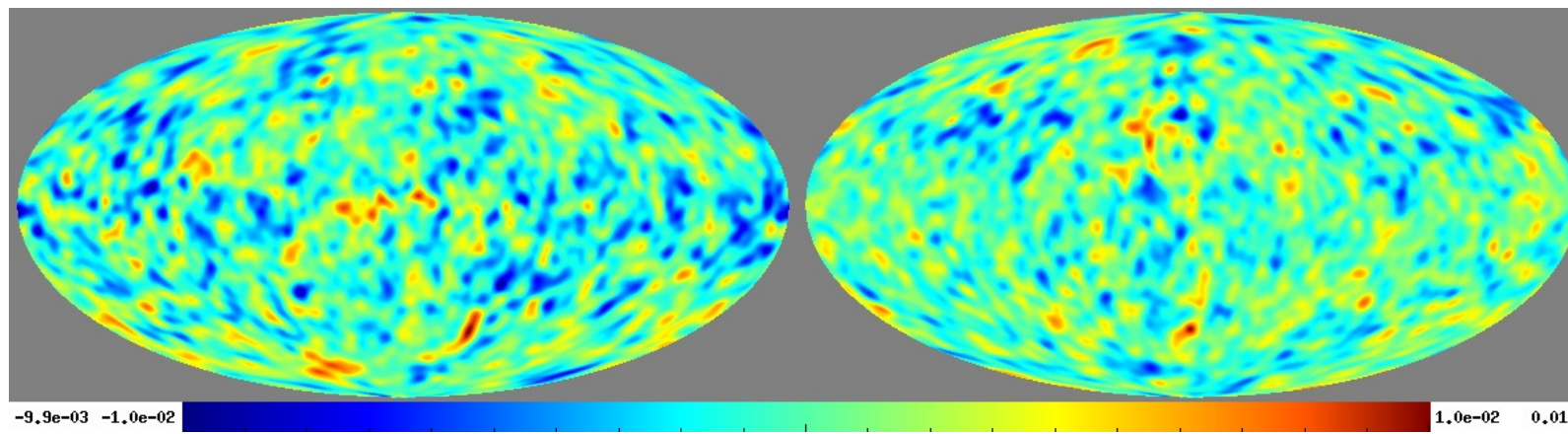
# CMB map



Smoothed to a  $5^\circ$  beam



Input – output CMB map  
smoothed to a  $5^\circ$  beam



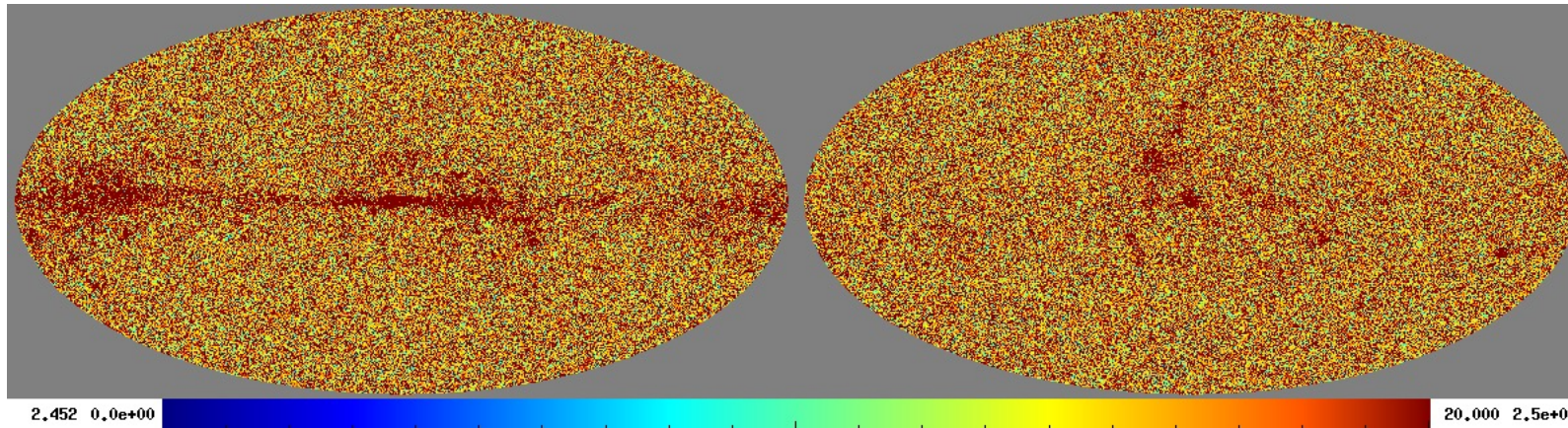
Q

$\mu K$

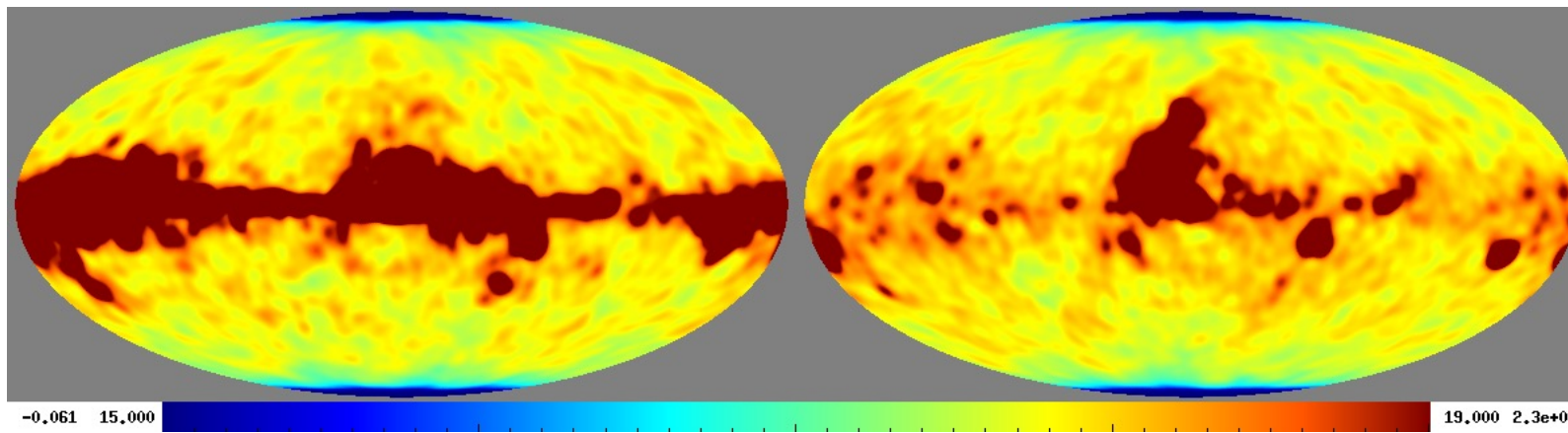
U

3

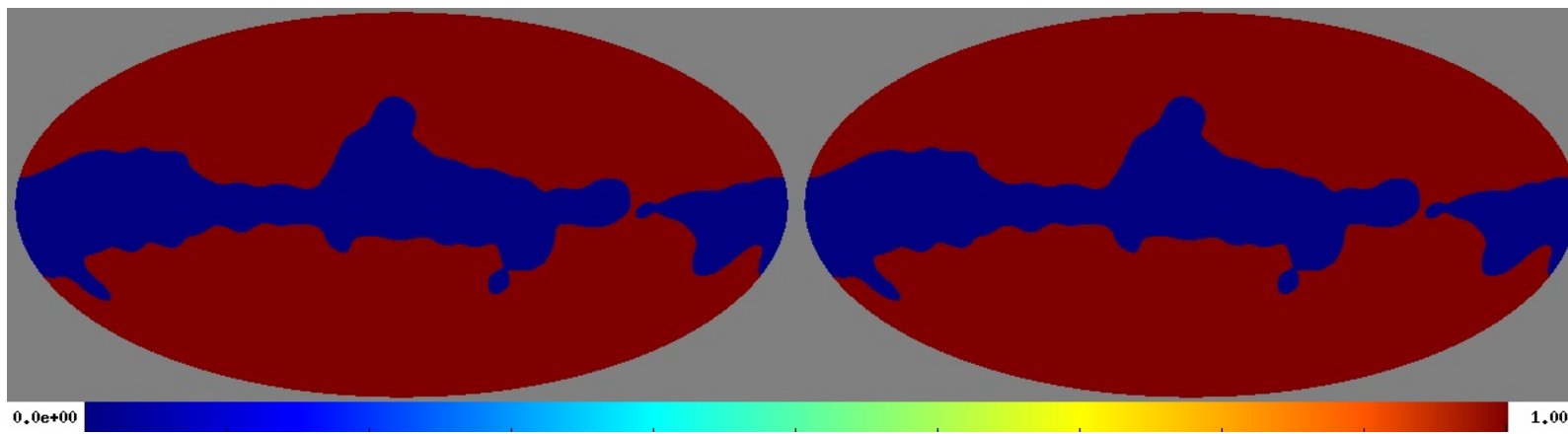
$\chi^2$  map



Smoothed to a 5° beam



25% mask



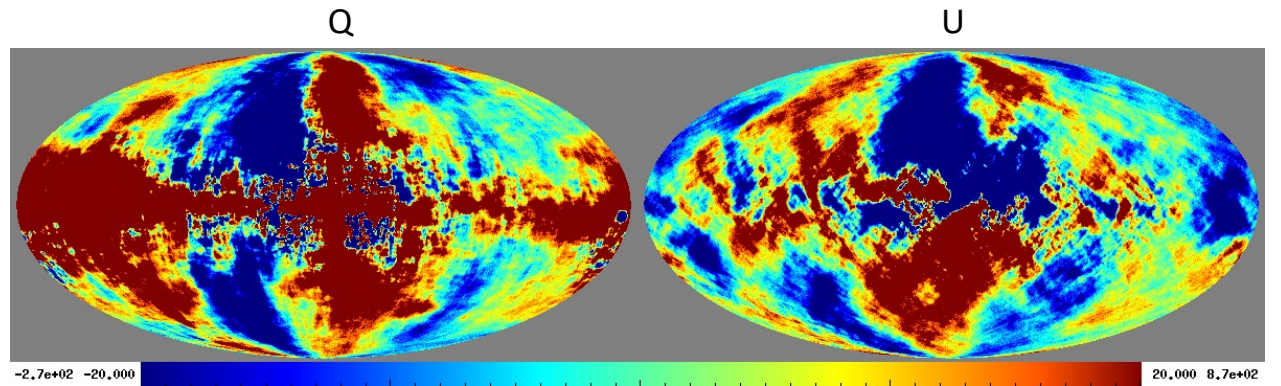
Q

U

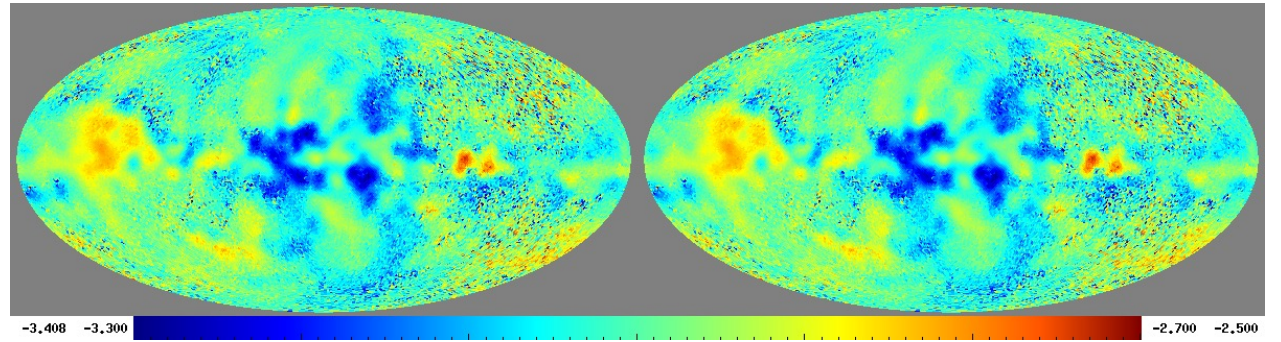
4

# Synchrotron maps

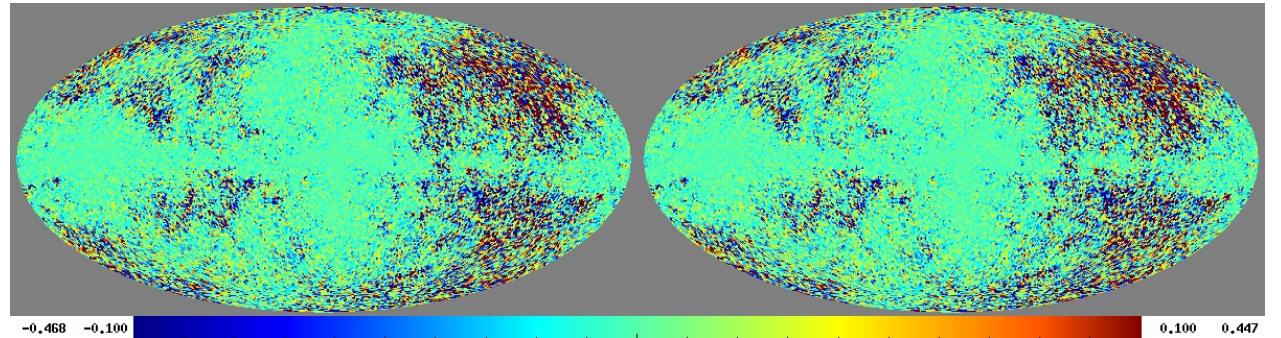
Amplitude



Spectral index,  $\beta_S$



Spectral index diff  
(input-output)

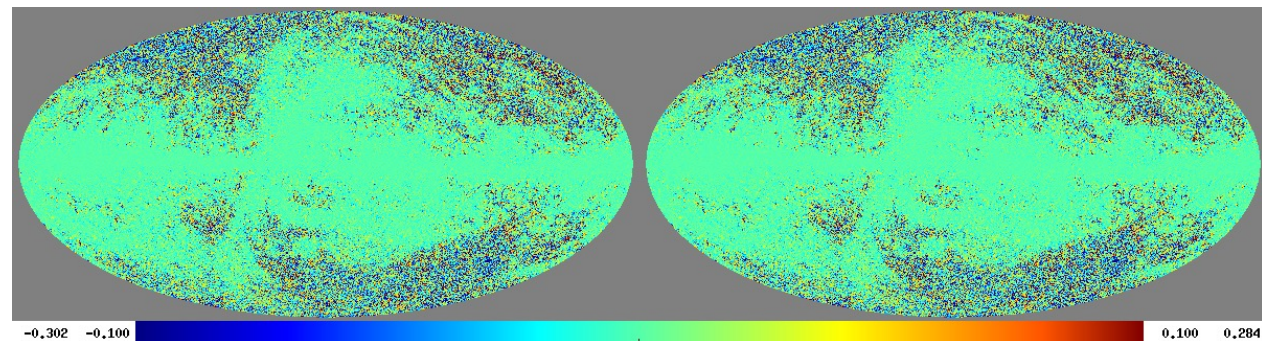
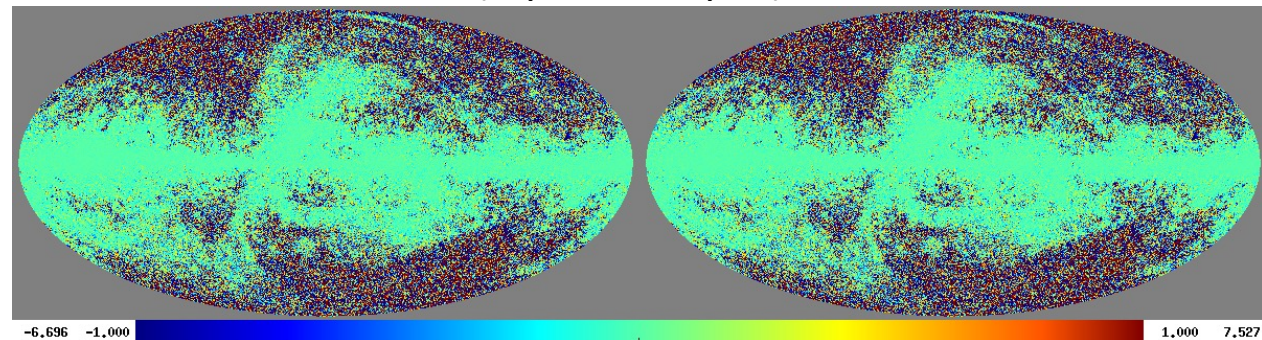


# Dust maps

Difference maps  
(input – output)

Q

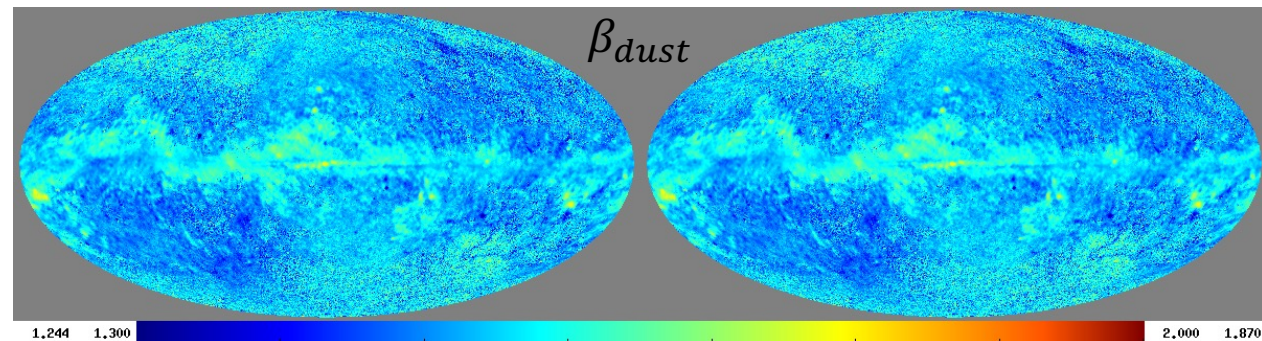
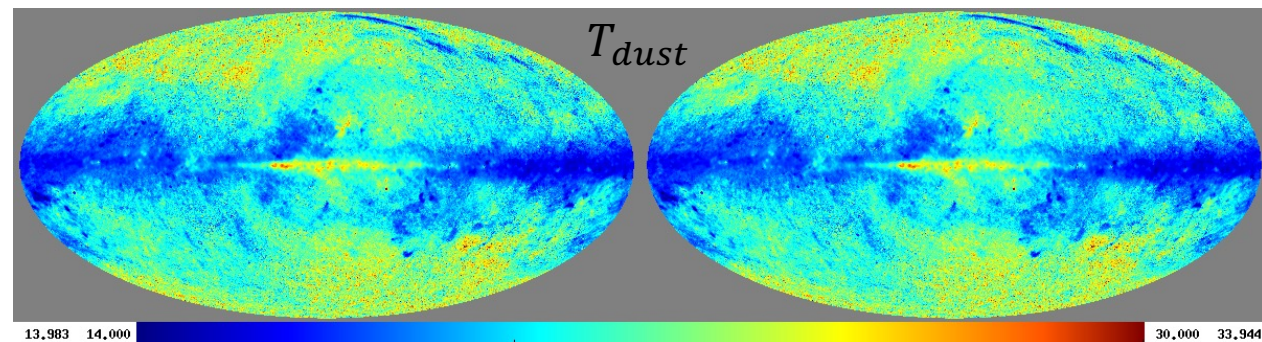
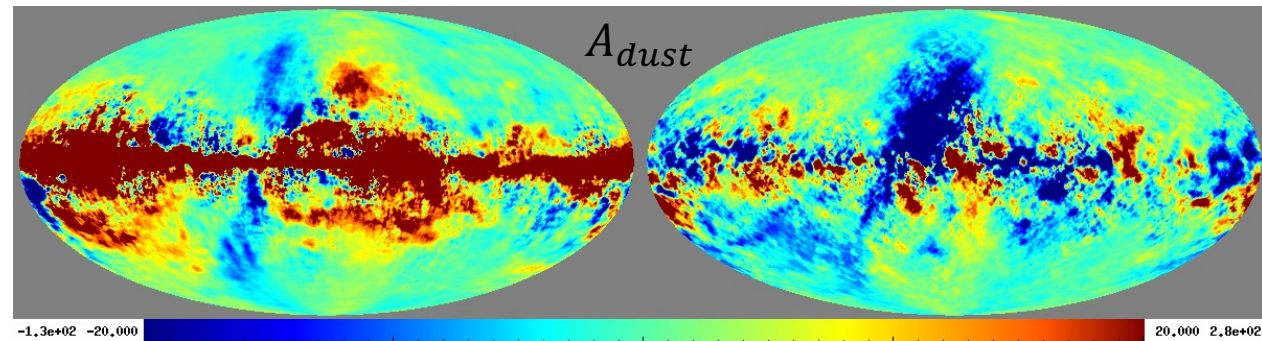
U



Output maps

Q

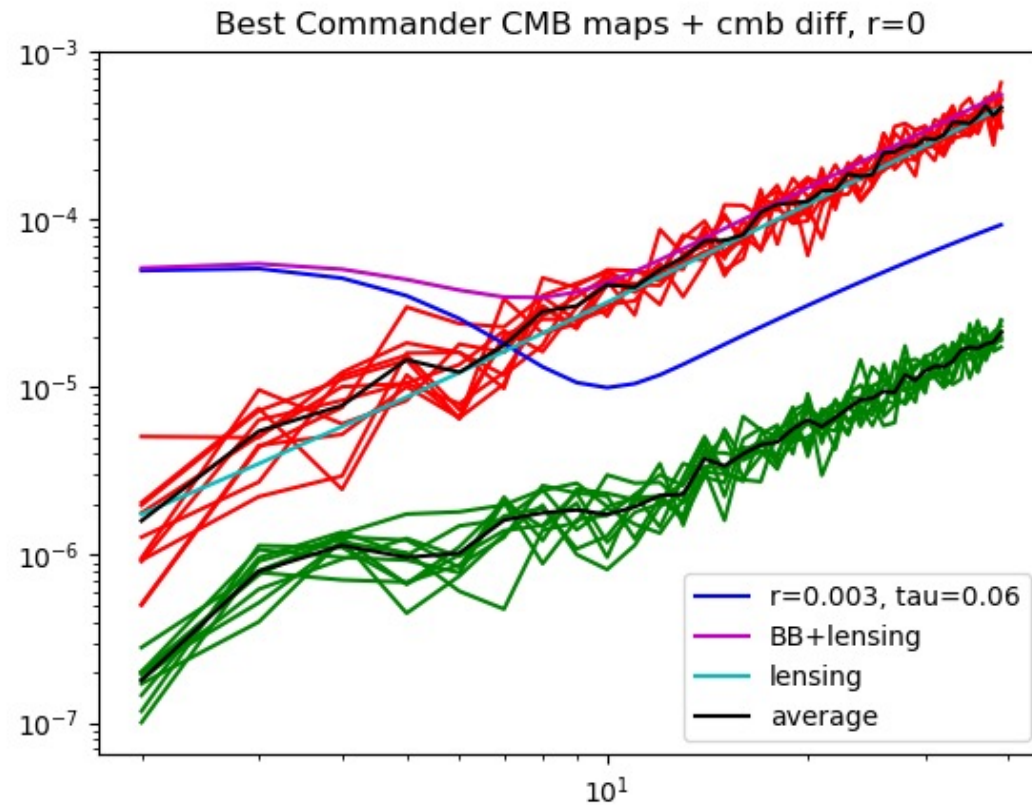
U



# Powerspectrum of CMB maps and CMB difference maps

$r = 0$

- Unmasked maps
- Best CMB map for all 10 simulation set (red lines)
- Subtract input and output CMB map (green lines)
- No masking or infilling

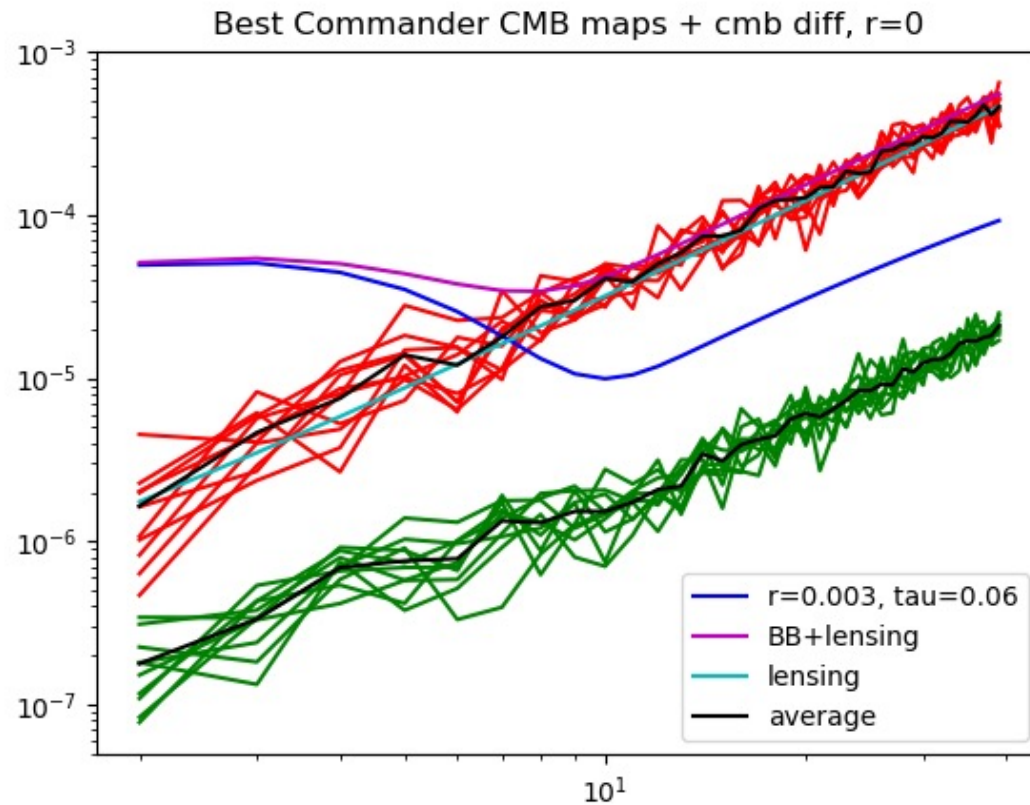


# Powerspectrum of CMB maps and CMB difference maps

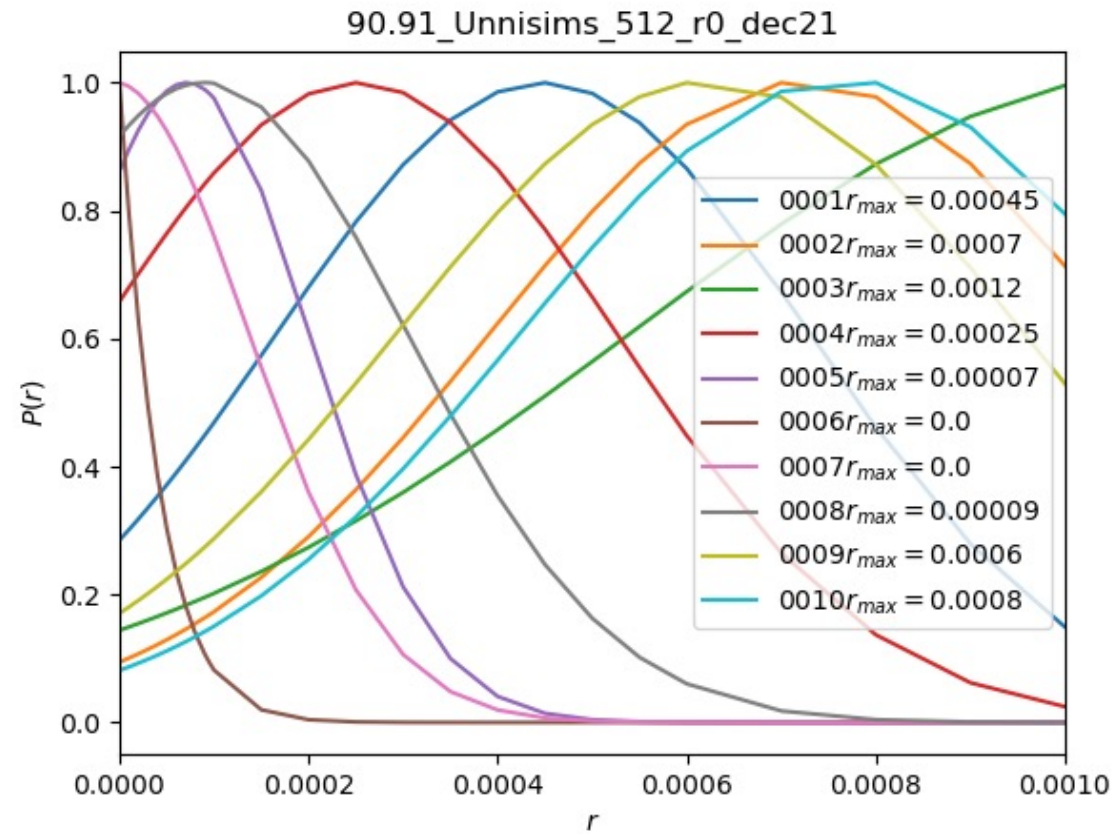
$$r = 0$$

New, improved dust priors

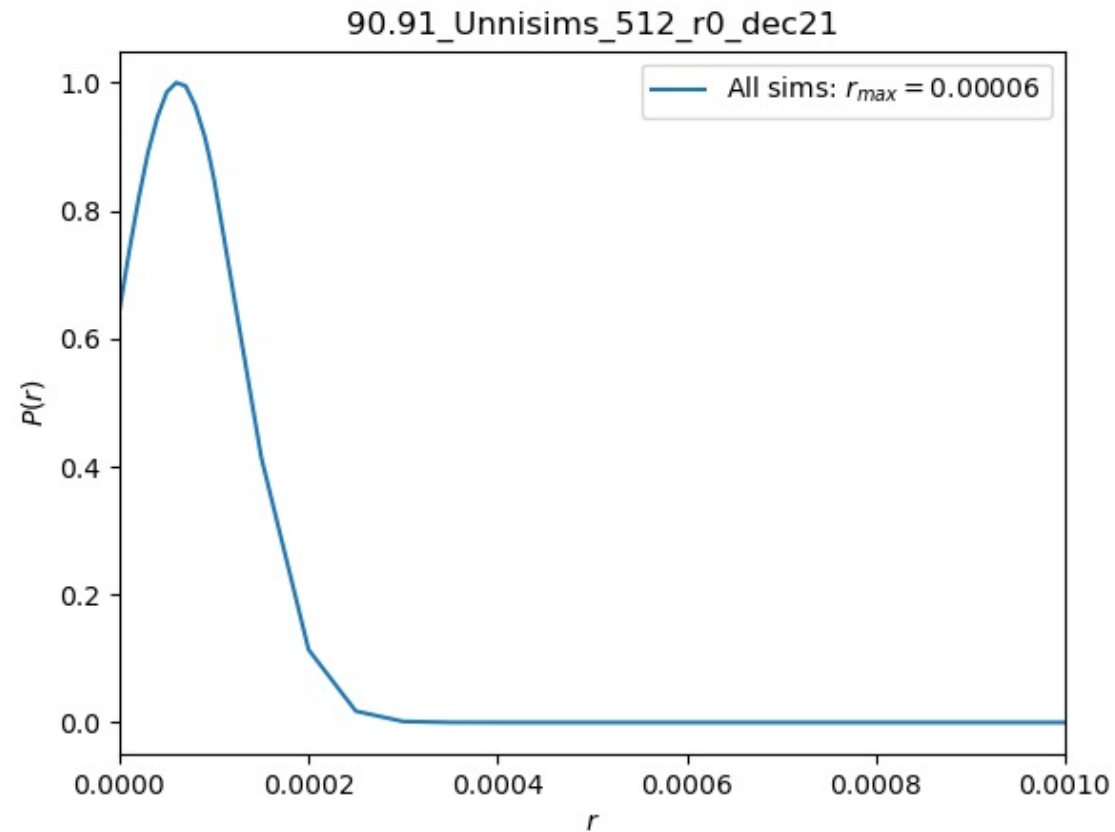
- Unmasked maps
- Best CMB map for all 10 simulation set (red lines)
- Subtract input and output CMB map (green lines)
- No masking or infilling



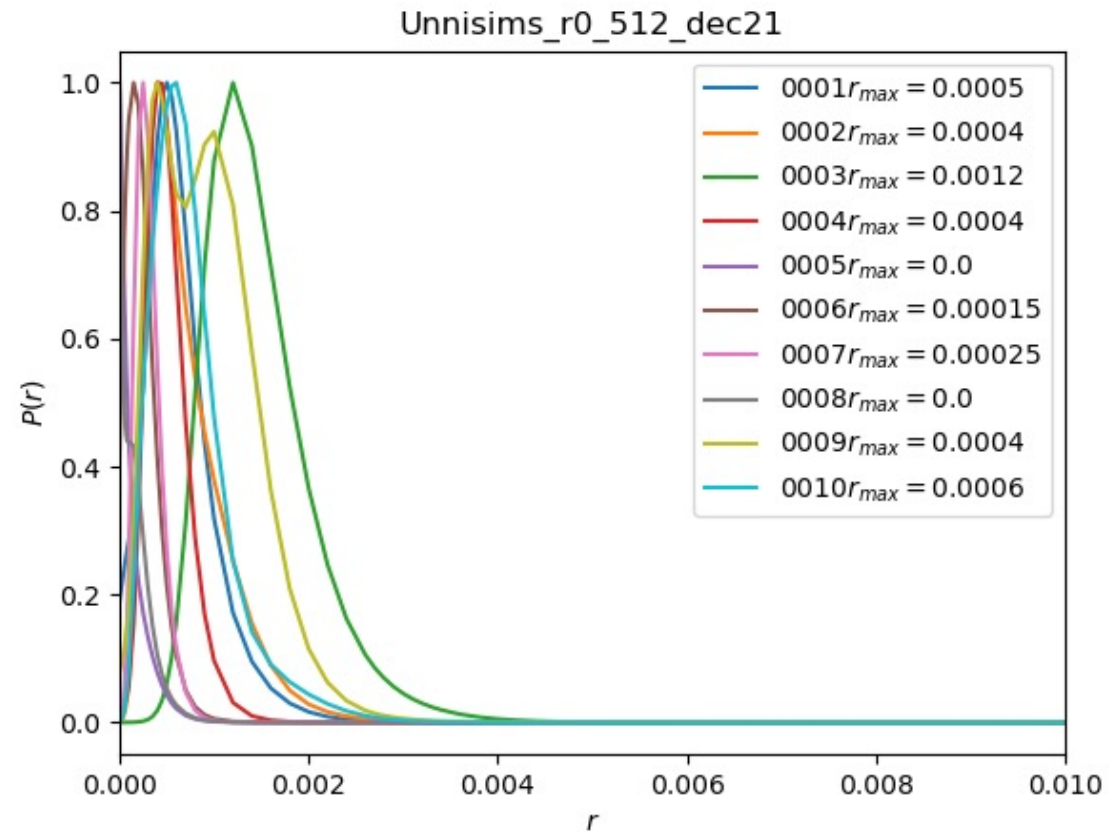
# Gaussian likelihood, $r = 0$



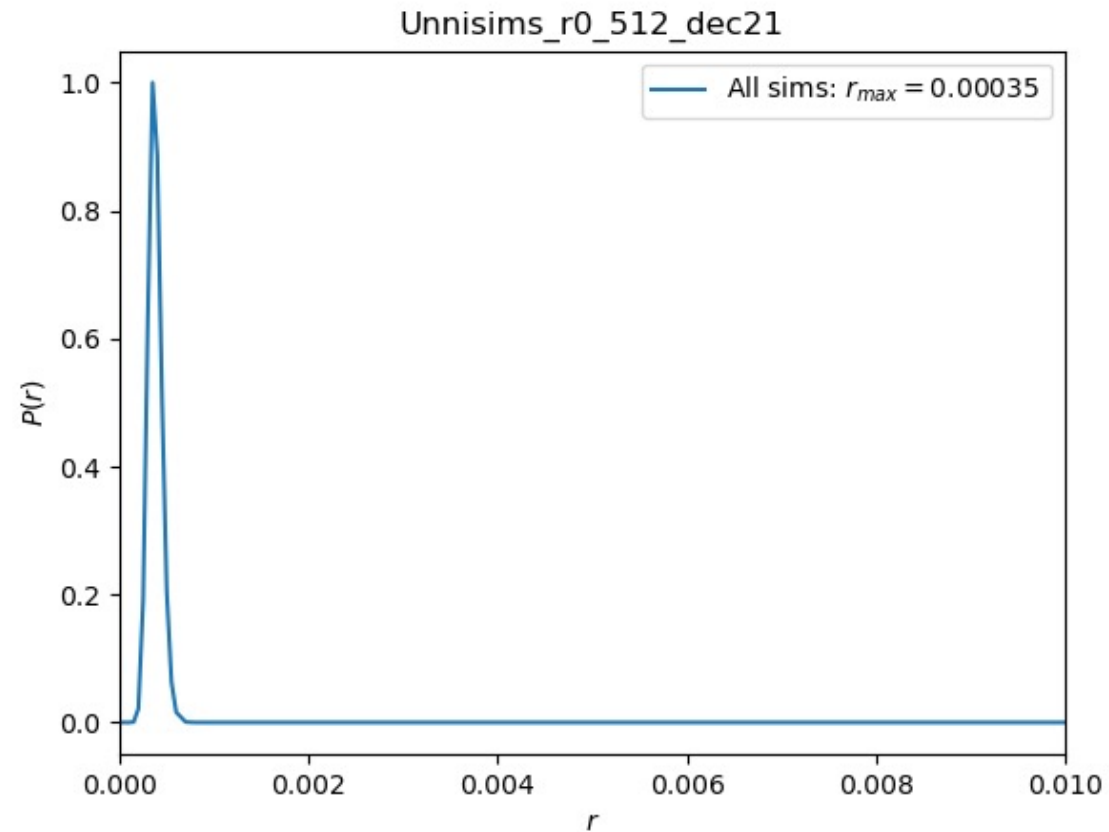
# Gaussian likelihood – total for all 10 realizations, $r = 0$



# Blackwell Rao estimation, $r = 0$



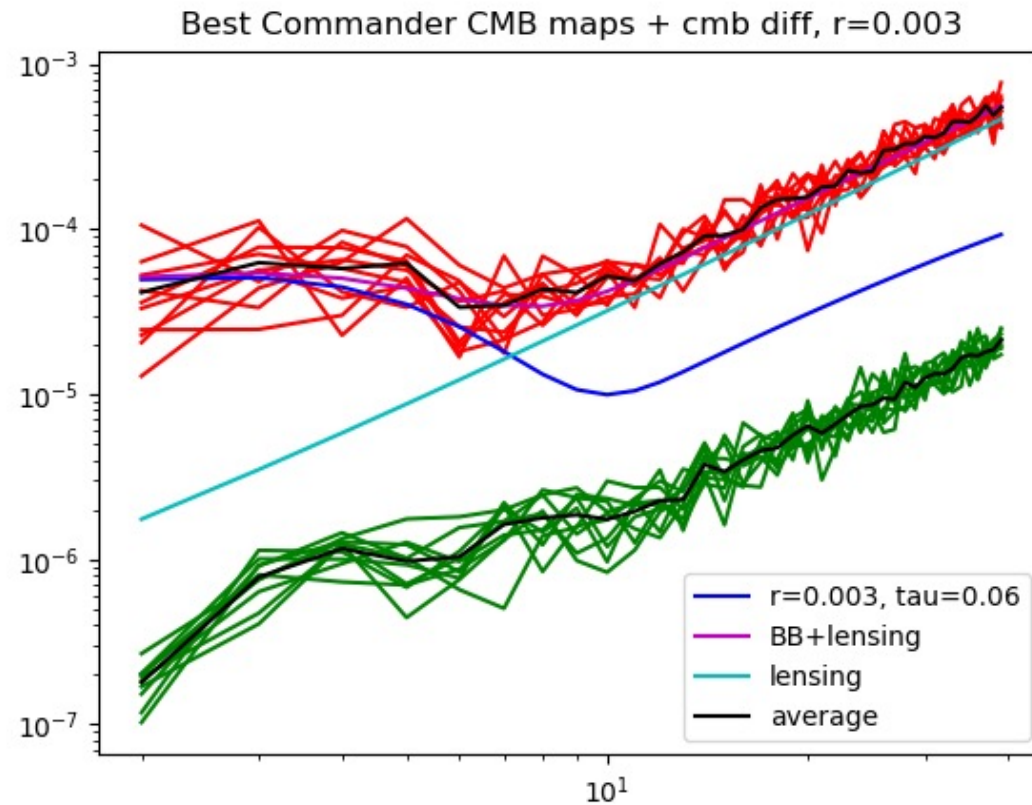
# Blackwell Rao estimation – total for all 10 realizations, $r = 0$



# Powerspectrum of CMB maps and CMB difference maps

$r = 0.003$

- Unmasked maps
- Best CMB map for all 10 simulation set (red lines)
- Subtract input and output CMB map (green lines)
- No masking or infilling

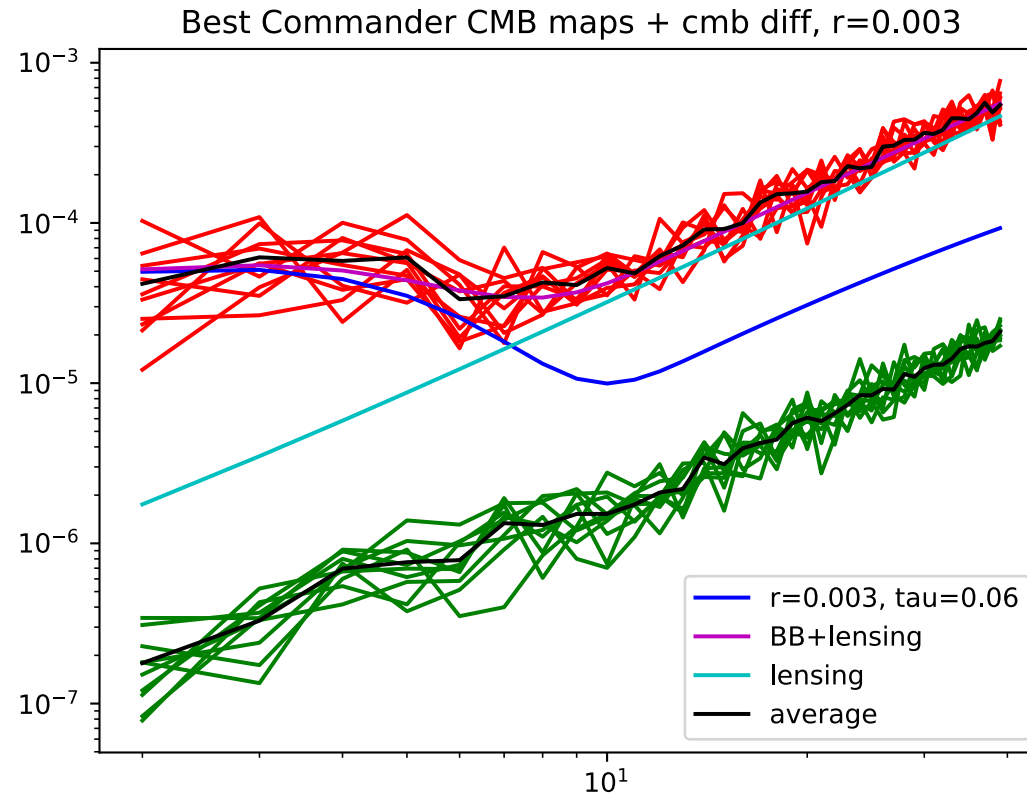


# Powerspectrum of CMB maps and CMB difference maps

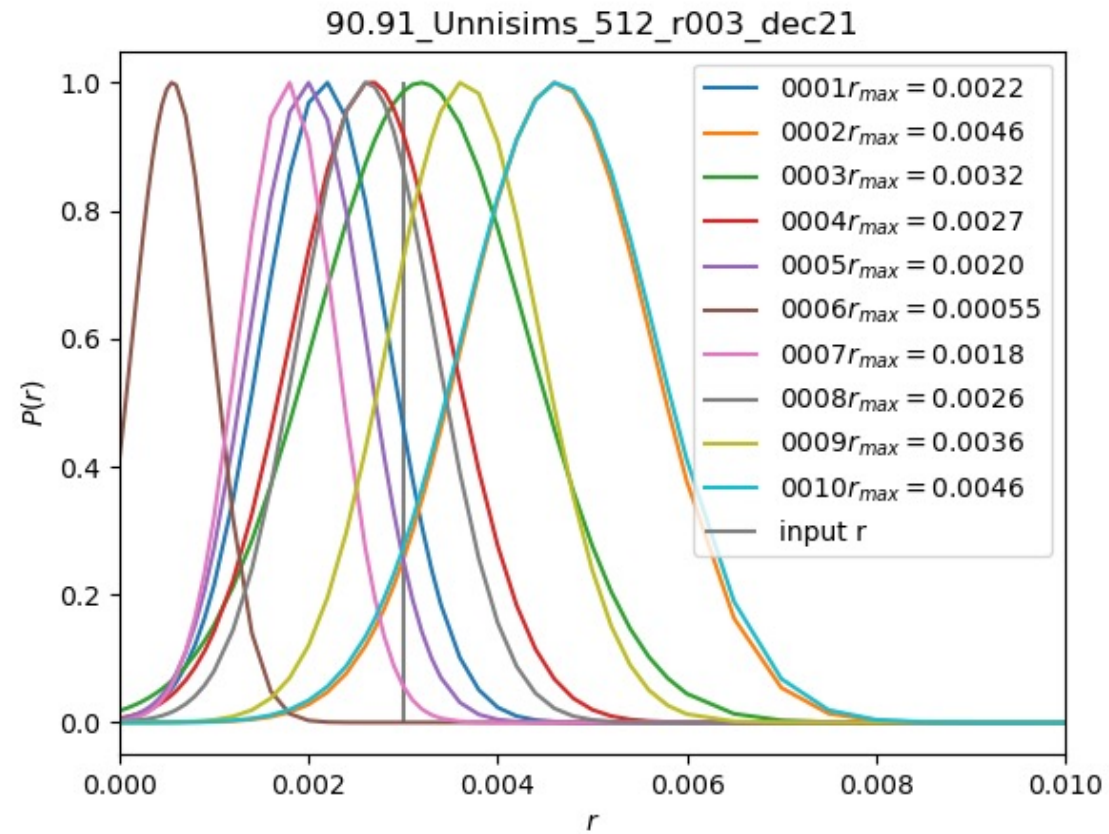
$$r = 0.003$$

New, improved dust priors

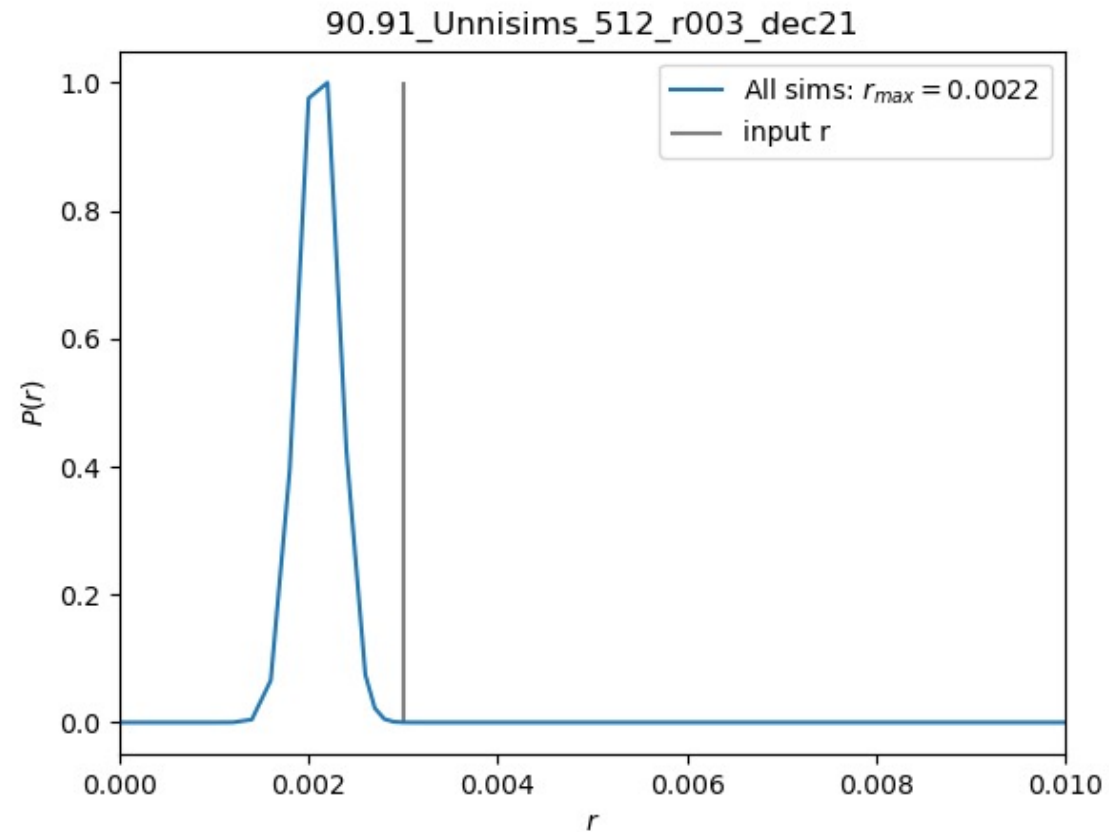
- Unmasked maps
- Best CMB map for all 10 simulation set (red lines)
- Subtract input and output CMB map (green lines)
- No masking or infilling



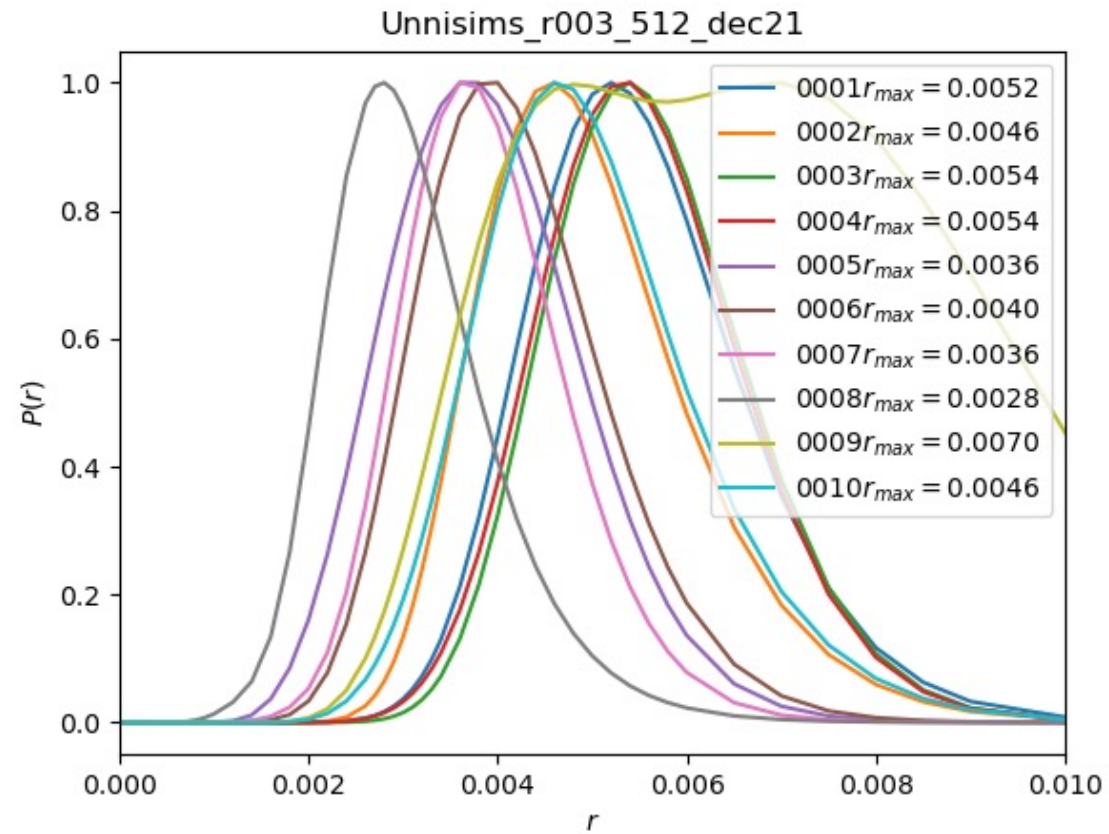
# Gaussian likelihood, $r = 0.003$



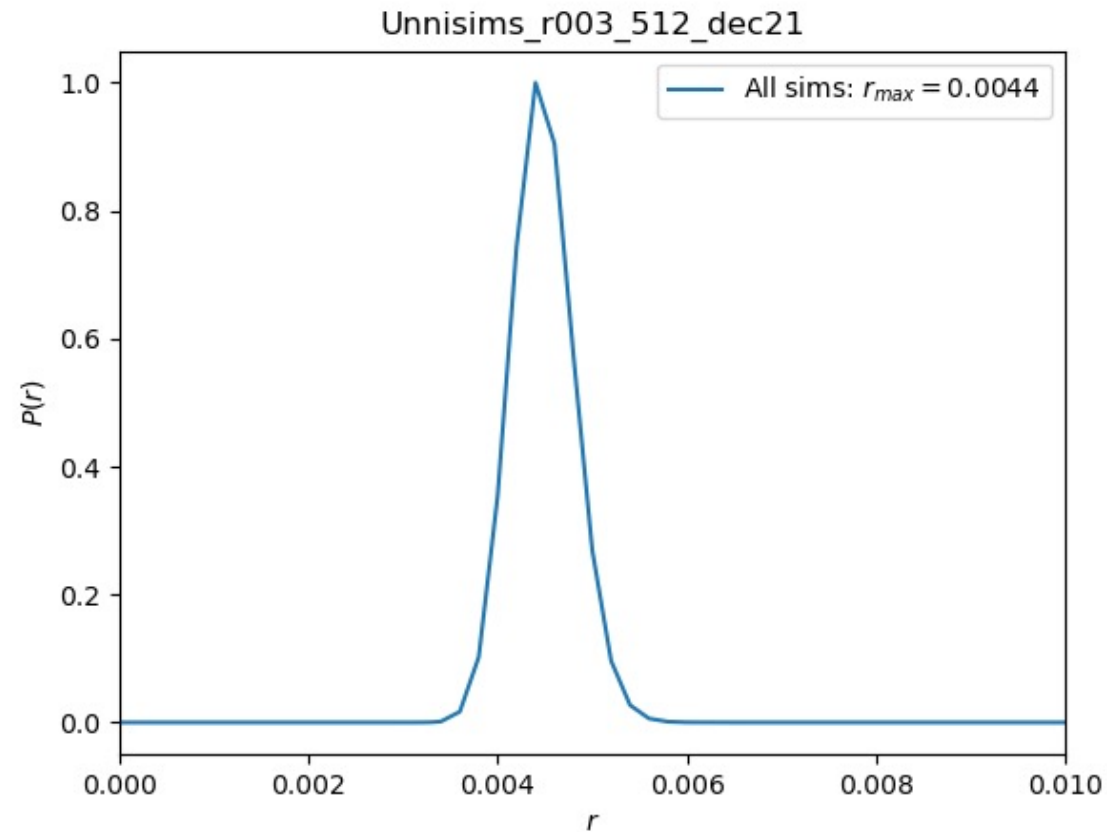
# Gaussian likelihood – total for all 10 realizations, $r = 0.003$



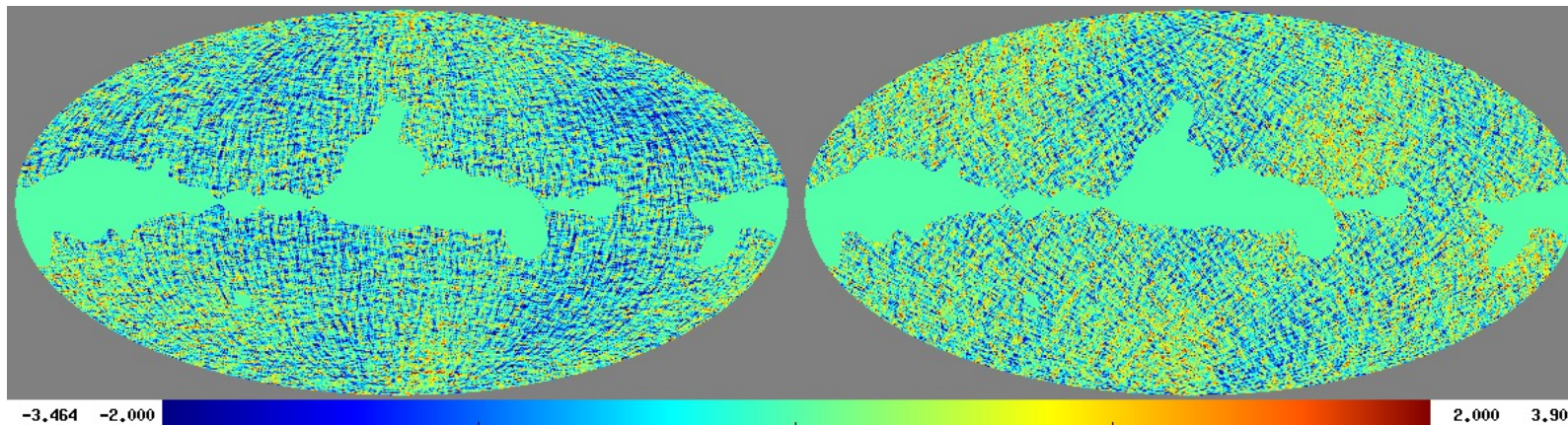
# Blackwell Rao estimation, $r = 0.0003$



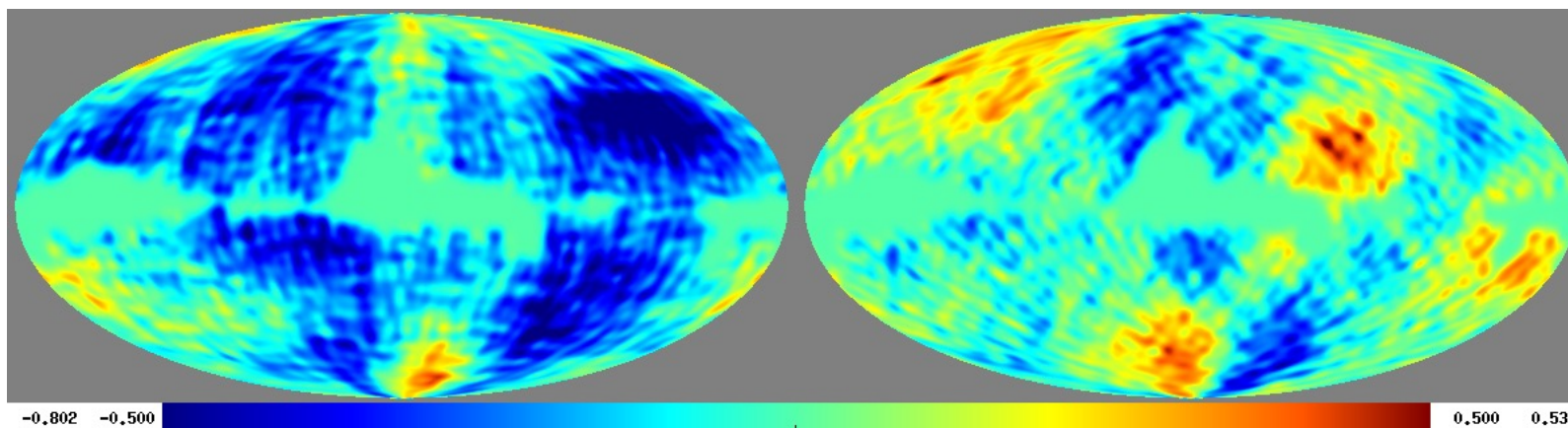
# Blackwell Rao estimation – total for all 10 realizations, $r = 0.003$



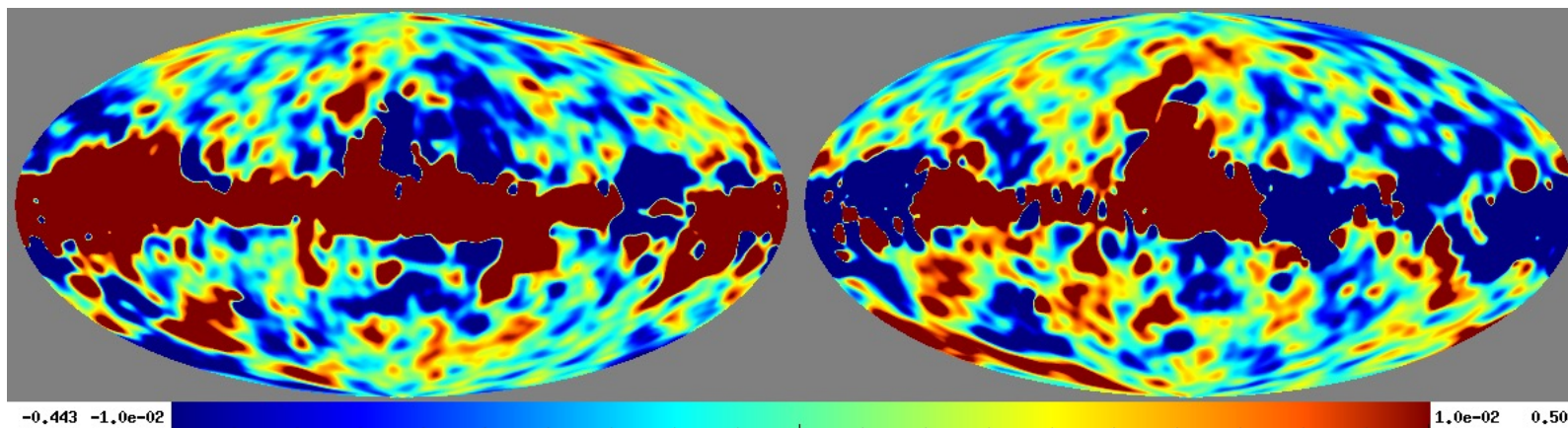
# 90.92 CMB



Smoothed to a 5° beam



Input – output CMB map  
smoothed to a 5° beam



Q

$\mu K$

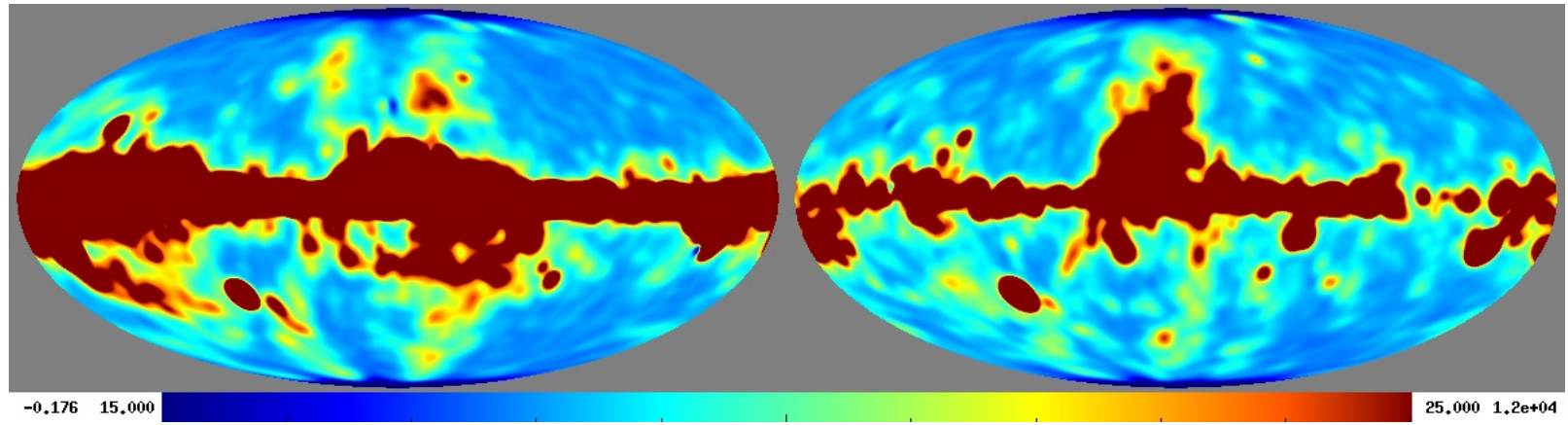
U

# Summary

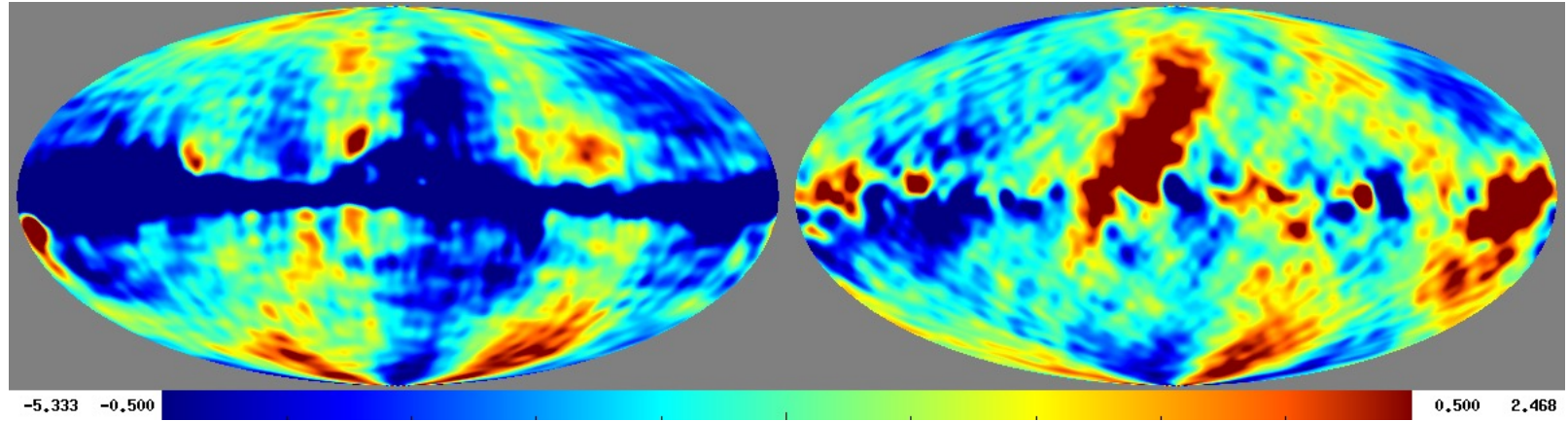
- Component separation gives results in excellent agreement with input
- Working on finding a likelihood that is both robust and unbiased
- Analysis of model 90.92 is not resulting in a clean enough CMB map

90.92 data  
with 90.91  
model

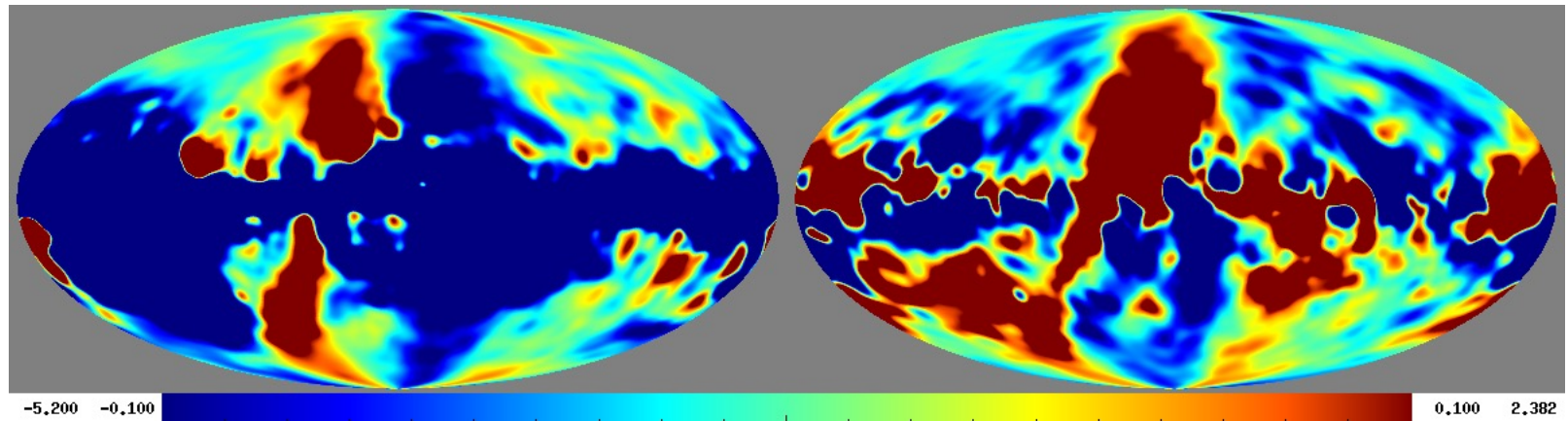
$\chi^2$  smoothed  
to a  $5^\circ$  beam



CMB smoothed  
to a  $5^\circ$  beam

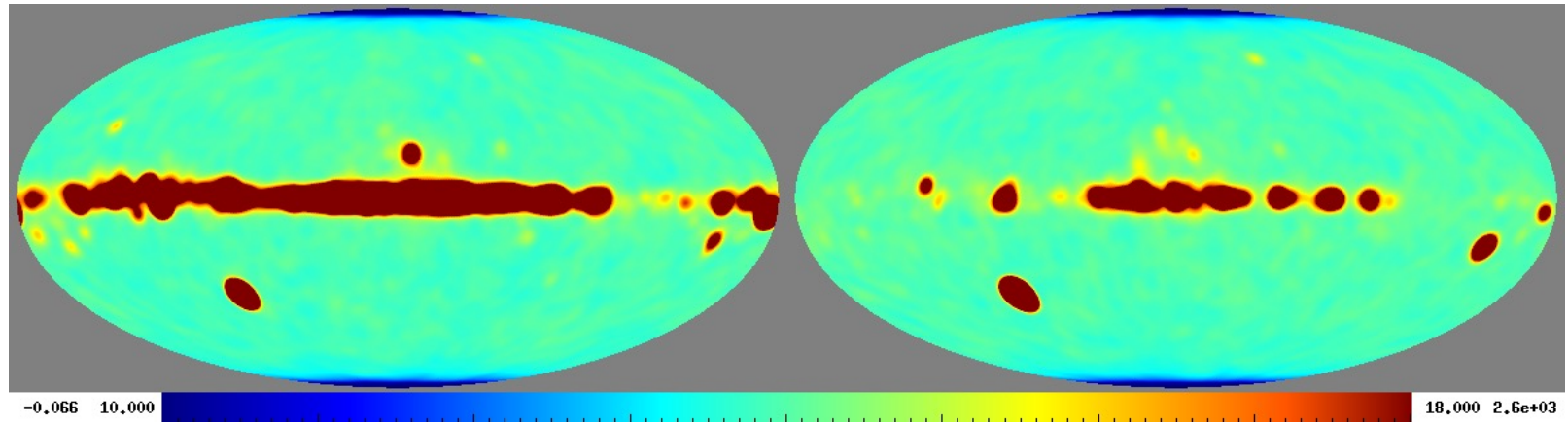


CMB diff  
(input - output)

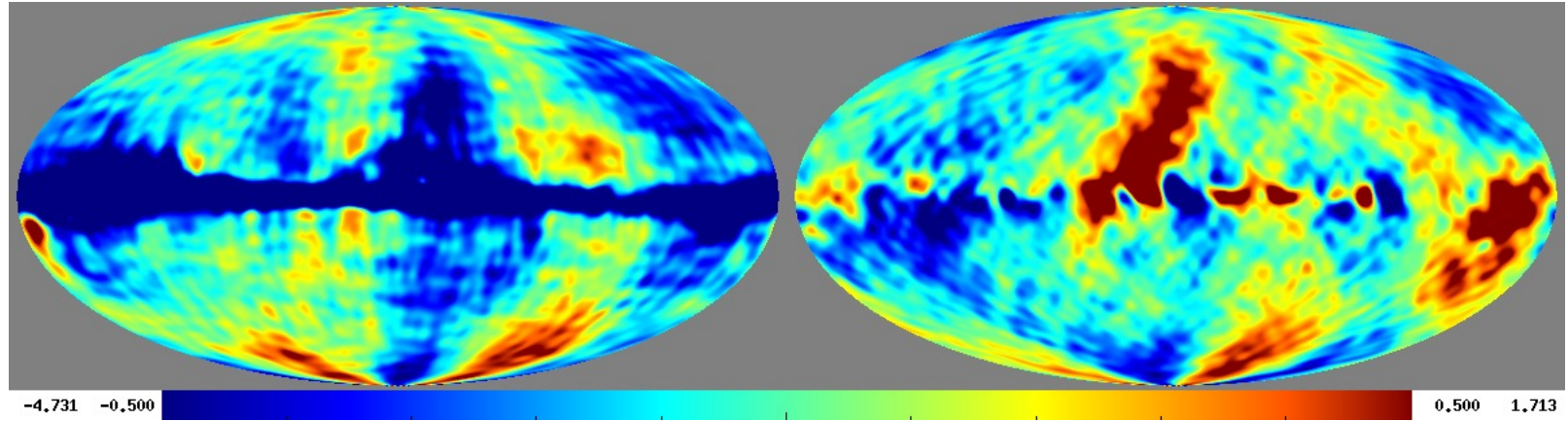


90.92 data  
with 90.91  
model  
21-385GHz

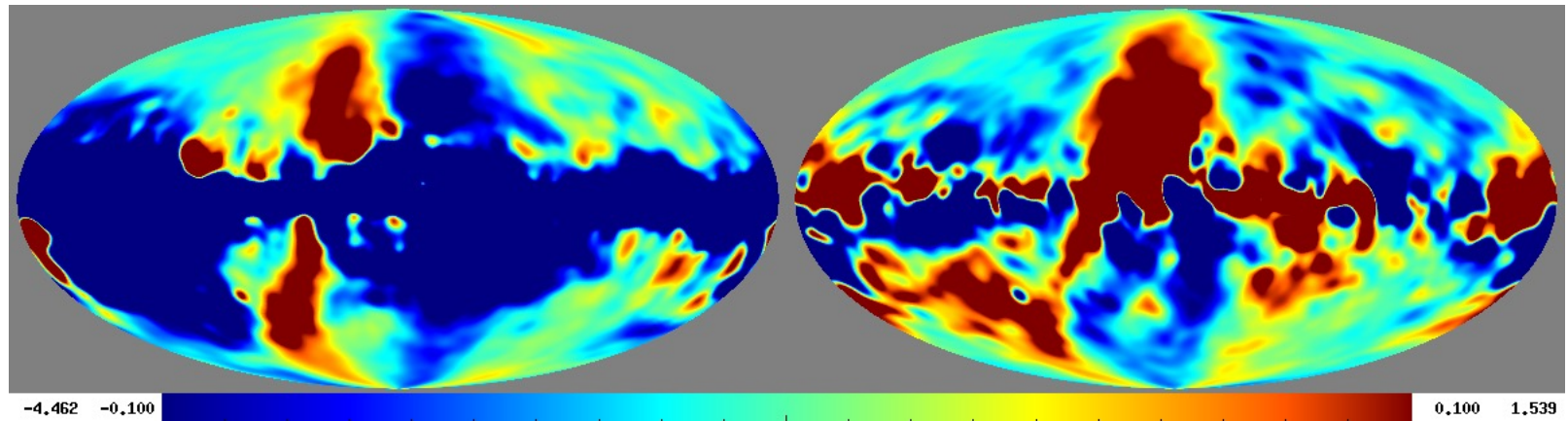
$\chi^2$  smoothed  
to a 5° beam



CMB smoothed  
to a 5° beam

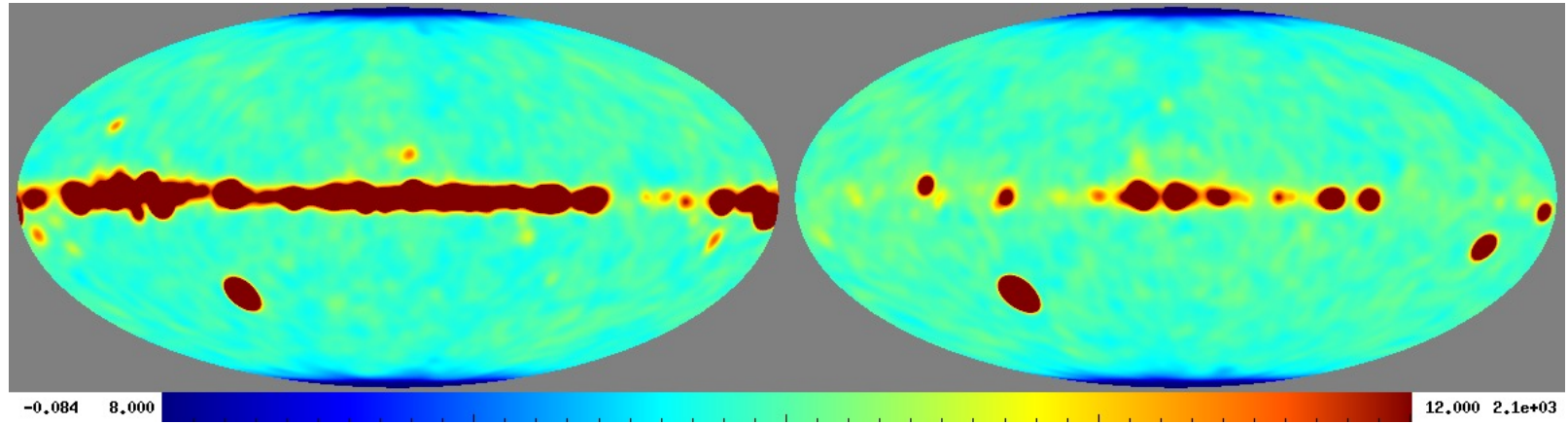


CMB diff  
(input - output)

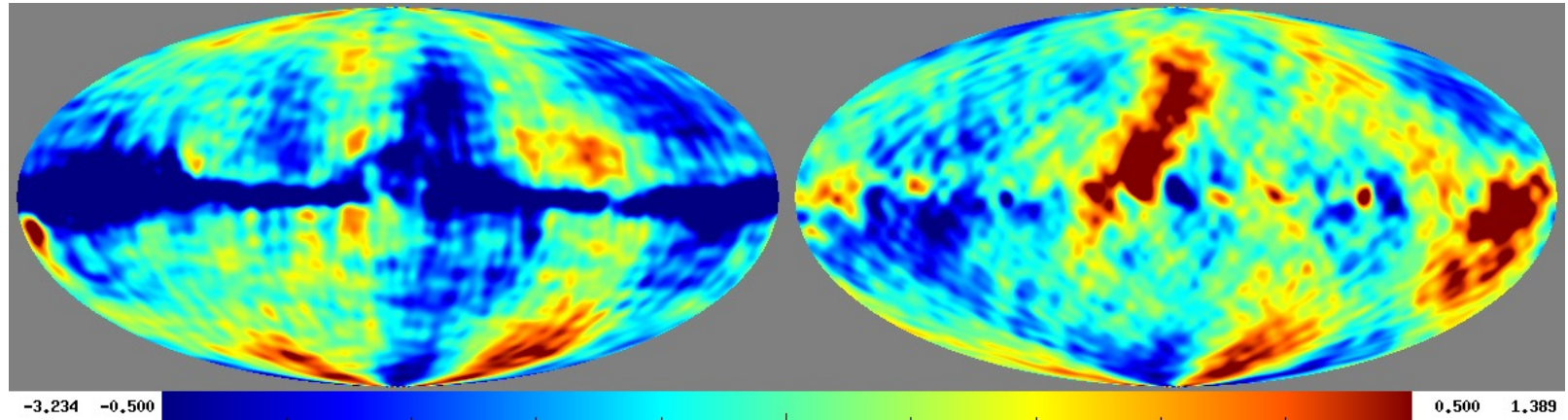


90.92 data  
with 90.91  
model  
43-385GHz

$\chi^2$  smoothed  
to a 5° beam



CMB smoothed  
to a 5° beam



Not visible in  
residuals

CMB diff  
(input – output)

